



Nonlinear Whistler Waves in Earth's Radiation Belts: THEMIS Observations

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Acknowledgements

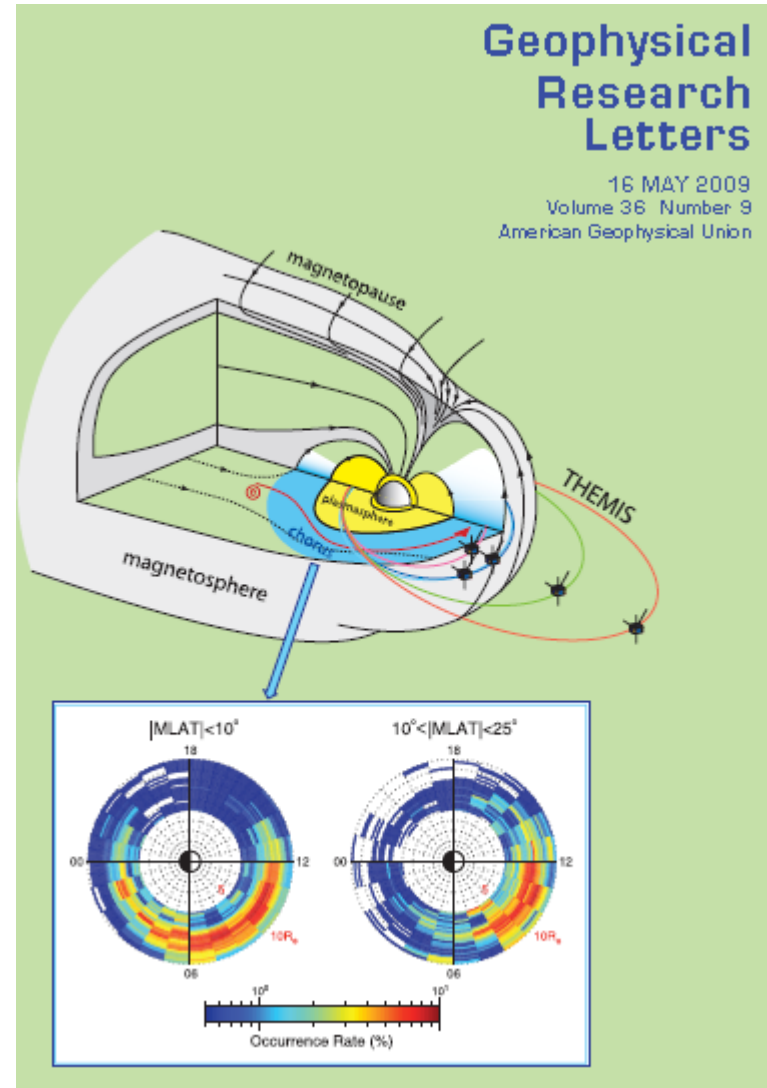
- Themis is the result of the work of many. In particular I'd like to acknowledge:
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 - ★ Electric Field Instrument:
 - John Bonnell, Forrest Mozer, Robert Ergun
 - ★ Particle Instruments:
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 - ★ Search Coil Magnetometer:
 - Alain Roux, Olivier LeContel
 - ★ Fluxgate Magnetometer:
 - Karl-Heinz Glassmeier

Outline

- Early observation:
 - ✦ Larger E amplitudes than expected
- Event study
 - ✦ Where quasilinear theory works
 - ✦ Where nonlinear theory is required
 - Chorus sweep
 - Omura's chorus generation mechanism
 - Experimental test
 - ✦ Cyclotron harmonic emissions

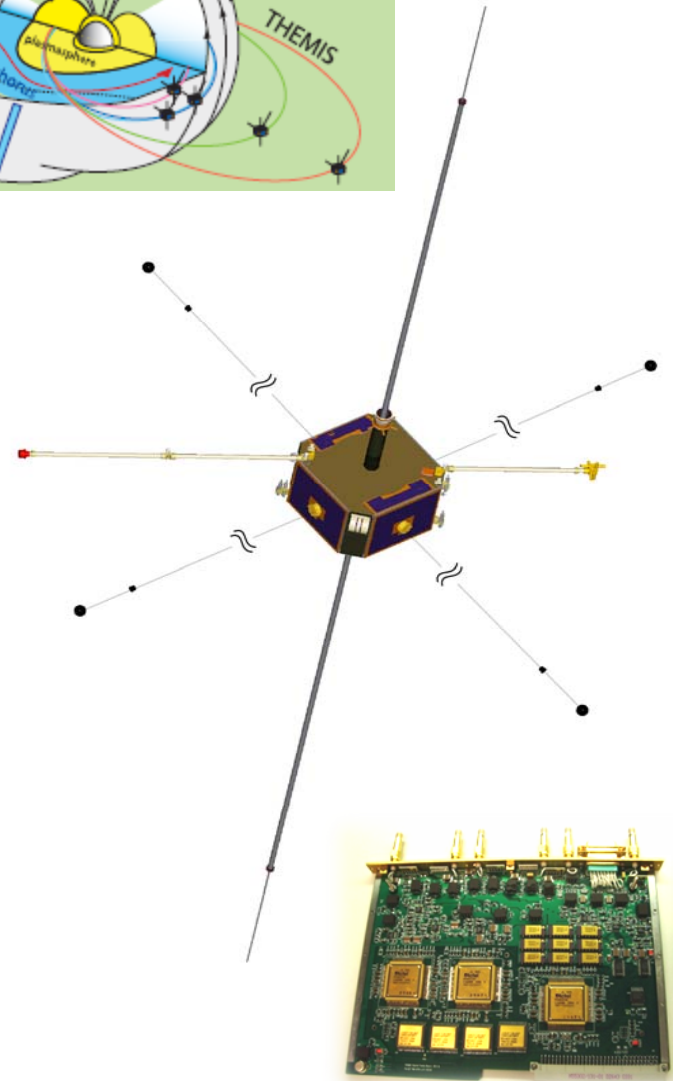
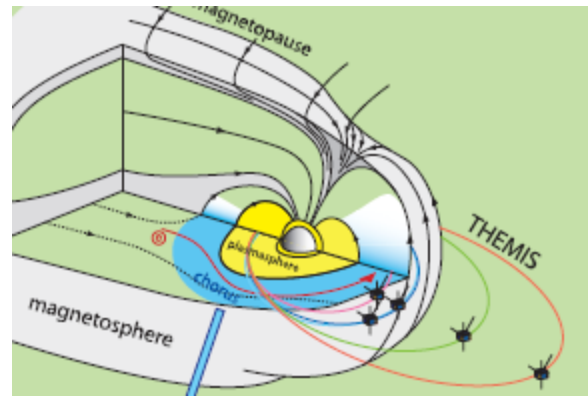
Background

- Wave-particle interaction is a major controlling process in the Earth's radiation belts
 - ★ Acceleration
 - ★ Loss (scattering into loss cone)
- Whistler waves
 - ★ Largely at dawn
 - ★ Equatorial generation
 - ★ Bursts of rising tones (chorus)



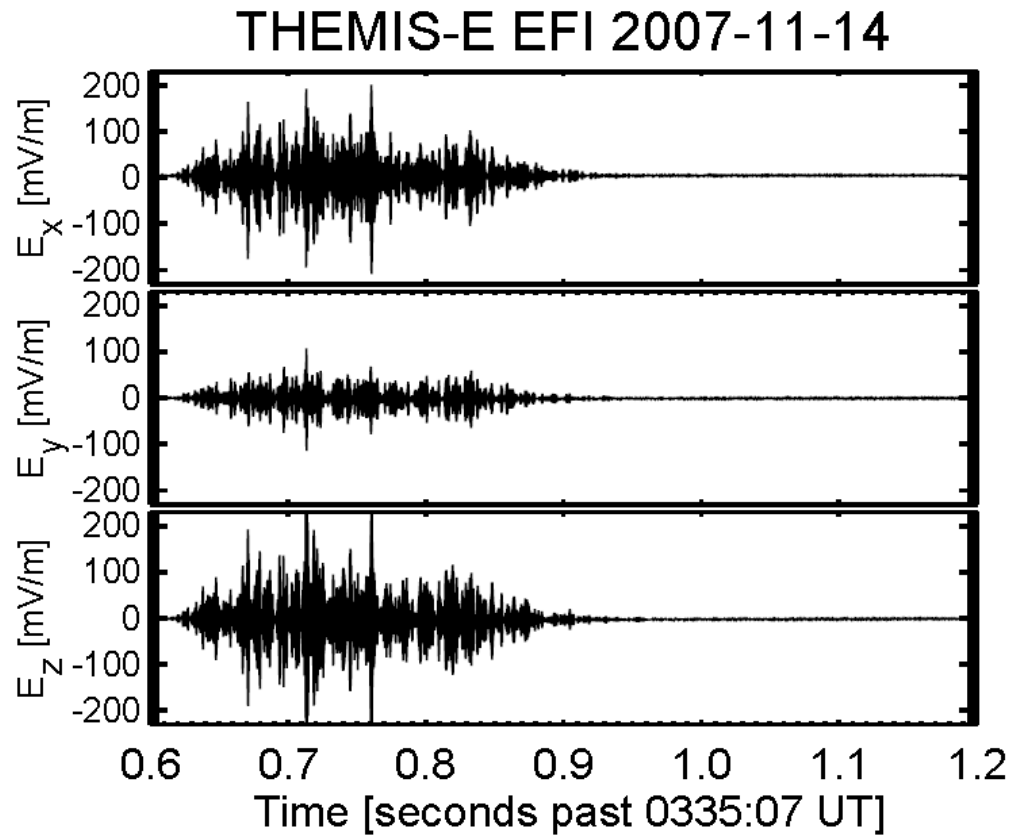
Themis

- 5 satellites in equatorial orbits
- Plasma instruments
 - ★ Low energy (<30 keV) + high energy (<6 MeV) ions and electrons
 - ★ Electric and magnetic fields



Unexpected amplitudes

- Early Themis result: Whistler electric fields larger than expected
- Heavy-tailed distribution
- >100 mV/m
 - ✦ Deterministic dynamics rather than stochastic
- Occur in large (hours of MLT) and persistent (days) region
- Agrees with other recent data (eg. STEREO)



Oblique propagation

- Many of the largest amplitude events are obliquely propagating
 - ✦ Even near the equator
 - ✦ Lack the characteristic chorus chirp
- Propagation effect?

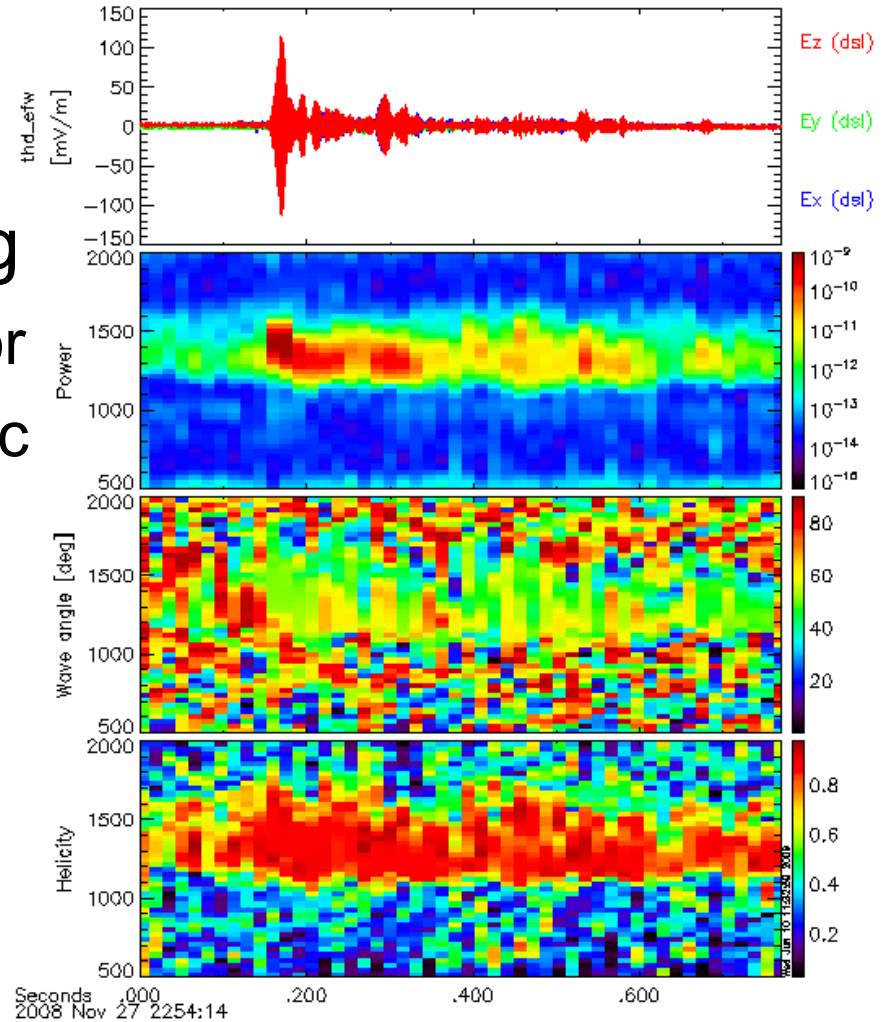
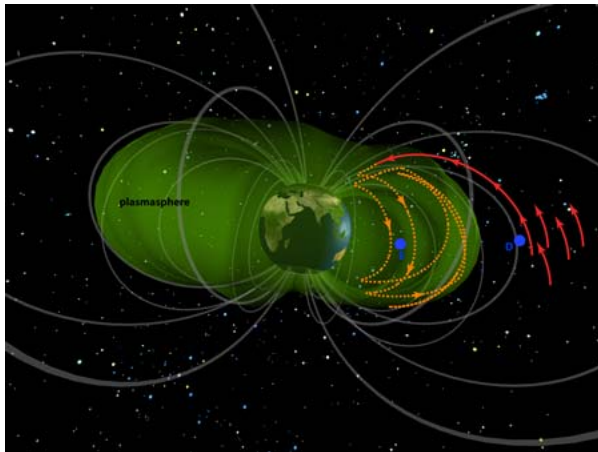


Image: NASA/Jacob Bortnik. See Bortnik et al, Science, 2009.



Themis: nonlinear whistlers

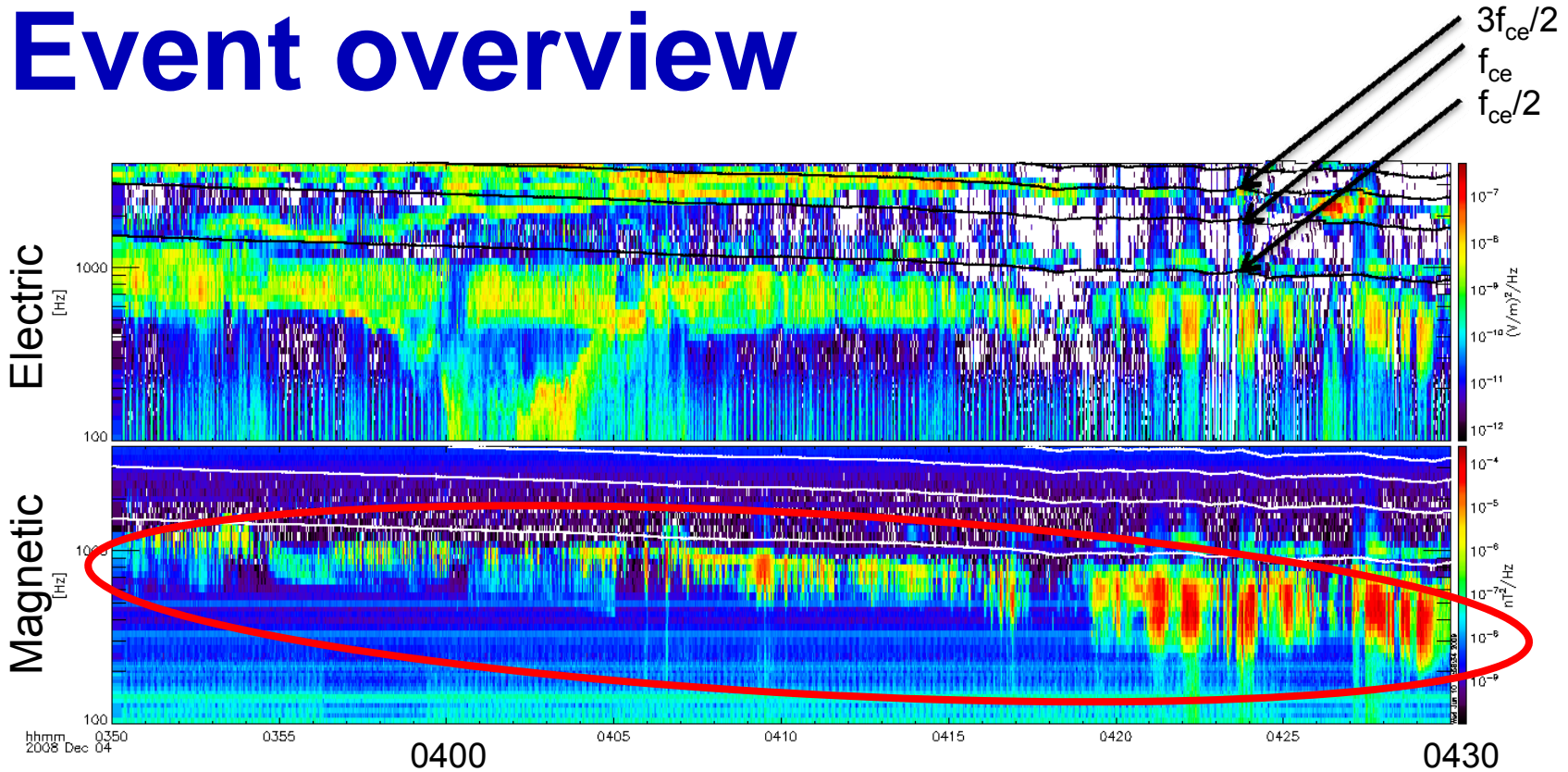
CASE STUDY



Outline

- Early result:
 - ✱ Larger E amplitudes than expected
- Event study
 - ✱ Where quasilinear theory works
 - ✱ Where nonlinear theory is required
 - Chorus sweep
 - Omura's chorus generation mechanism
 - Experimental test
 - ✱ Cyclotron harmonic emissions

Event overview



- Themis-D, $L \sim 5 R_E$, $MLAT \sim 2^\circ$
- Bursts of whistler activity up to $0.5 f_{ce}$
 - ★ Gap at $0.5 f_{ce}$ (common)
 - ★ Parallel propagation
- Strong cyclotron harmonics

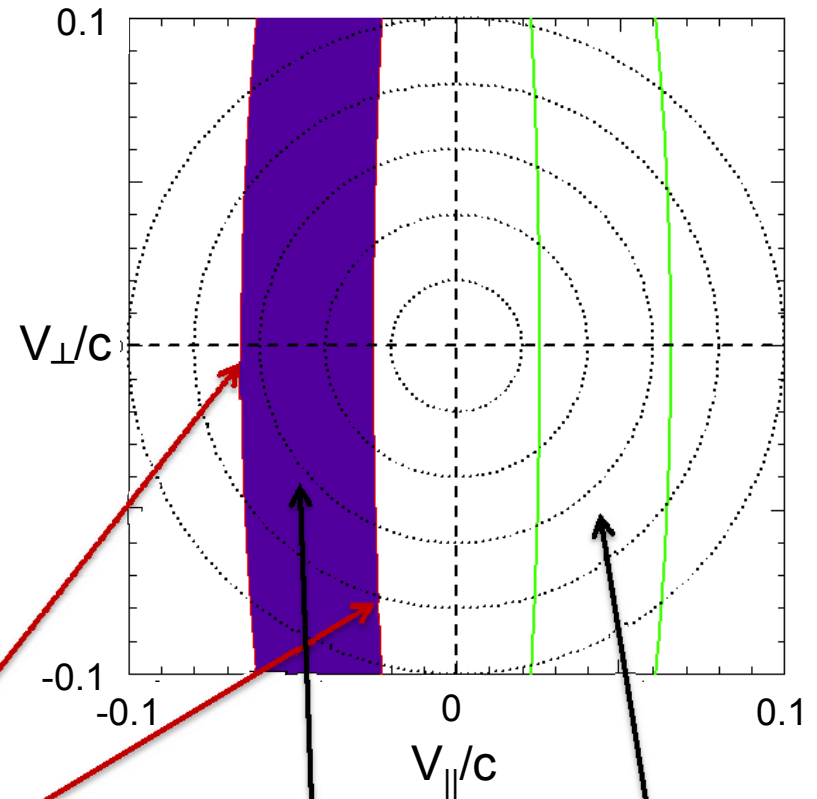
Resonant ellipses

- Electrons moving counter to the wave motion at

$$V_R = \frac{\omega}{k} \left(1 - \frac{\omega_{ce}}{\gamma\omega} \right)$$

are cyclotron-resonant with the wave

- Ellipses at $f=0.2 f_{ce}$ and $f=0.5 f_{ce}$ demarcate interaction boundaries

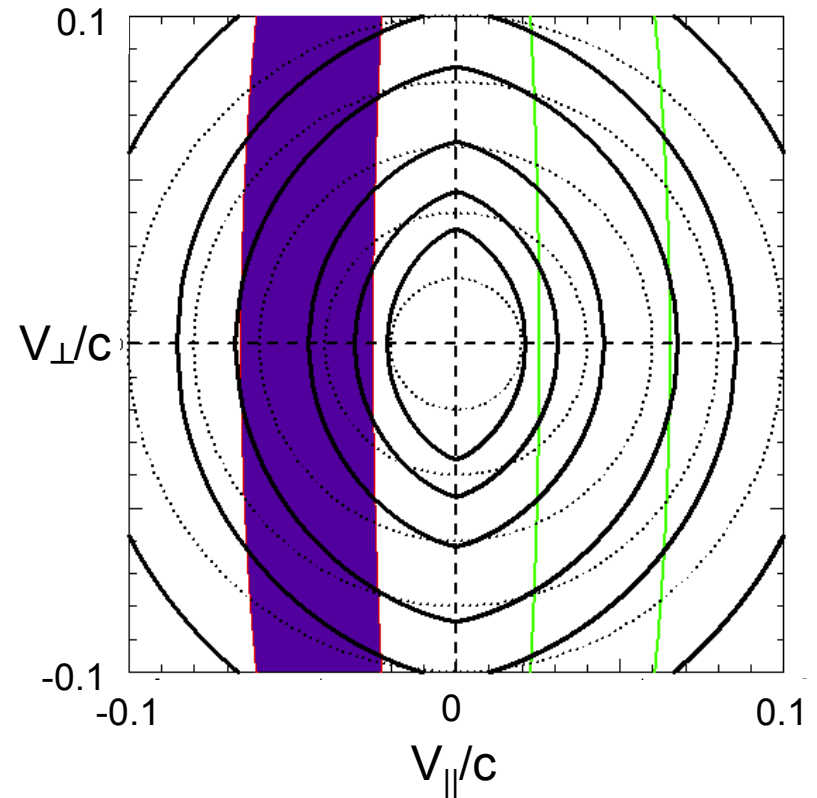


Region with
potentially
resonant
particles

Mirrored

Diffusion curves

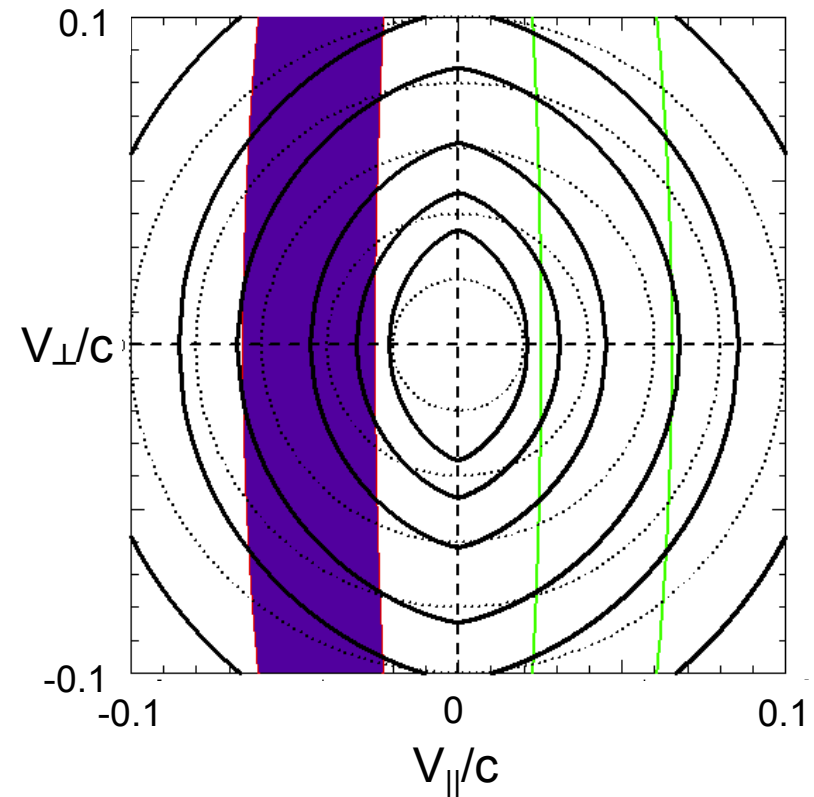
- Given broad spectrum of waves:
 - ✦ In frame moving with wave phase velocity, (relativistic) energy of resonant particle is conserved
 - ✦ As energy changes, particle resonates with different frequencies
- Result: diffusion curves shown
 - Hot, relativistic curves shown; cold and non-relativistic very similar



Summers, Thorne, Xiao, JGR, 1998
 Horne and Thorne, GRL, 2003
 Gendrin and Roux, JGR, 1980

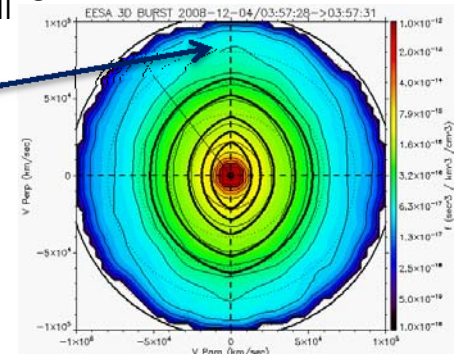
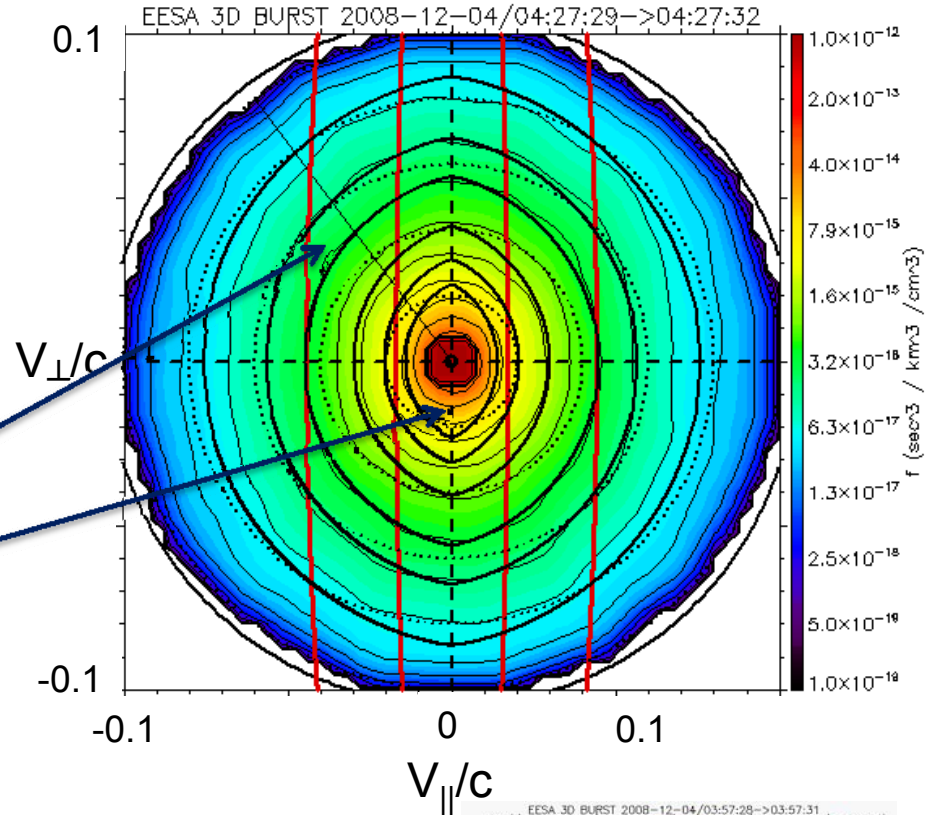
Experimental test

- Diffusion surface = phase space density iso-surface
 - Valid between the resonant ellipses for the observed wave spectrum
 - Marginal stability criterion



Diffusion curves: result

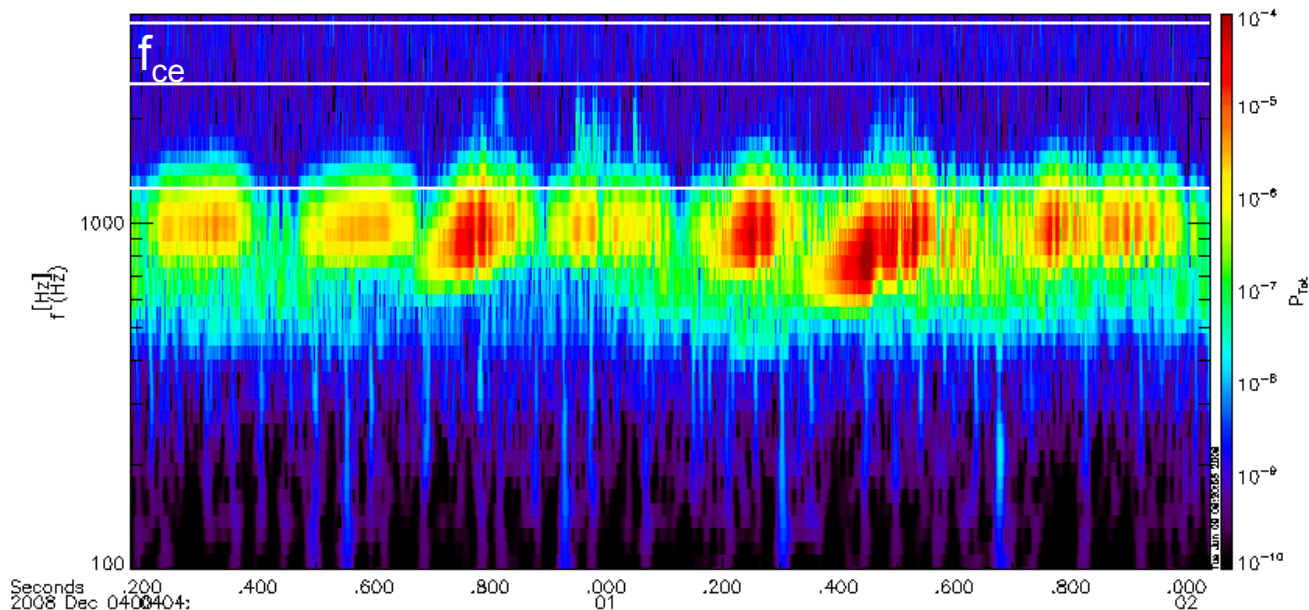
- Very good agreement with observed distribution function
- Diffusion curve=f iso-surface in interaction region
- Isotropic core
 - ✦ Resonant frequency above $f_{ce}/2$ (no wave power available)
- Sometimes see shoulder at $f_{ce}/2$



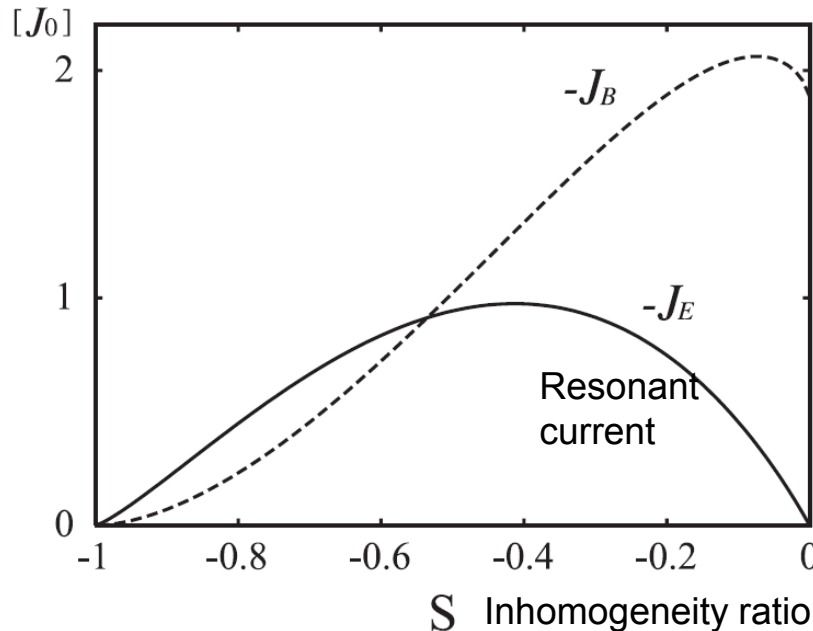
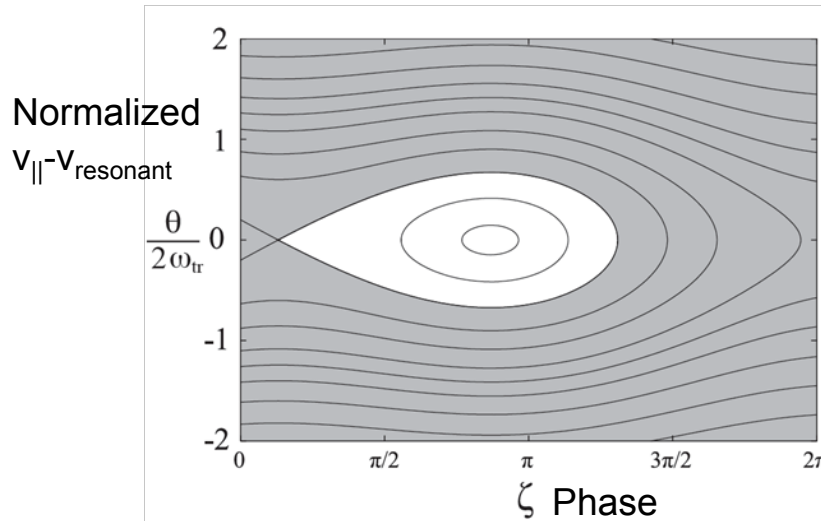
Nonlinear questions

Linear theory predicts wave growth in a broad band up to f_{ce} . So:

1. Why does the chorus element chirp?
2. Why is the wave absorbed at $f_{ce}/2$?

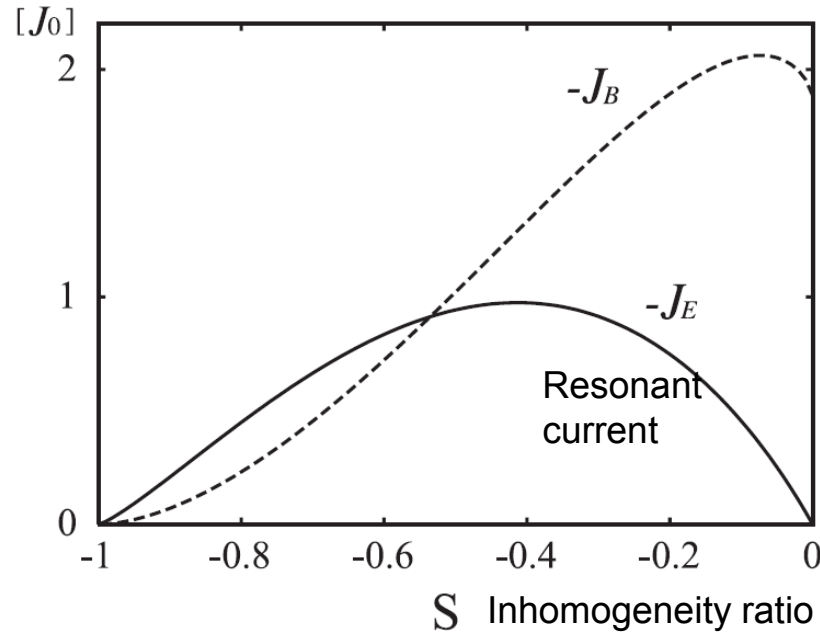


Growth by nonlinear trapping



- Omura, Katoh and Summers, JGR, 2008
- Wave trapping results in electromagnetic phase space hole
- For specific conditions, resonant currents cause wave growth
 - ★ Rising frequency
 - ★ Quantified by inhomogeneity ratio S

Omura et al.'s prediction



- Fastest-growing wave is the one that maximizes the resonant current
- ★ Inhomogeneity ratio $S \sim -0.4$

$$S = -\frac{m_0}{kv_{\perp} e B_w \delta^2} \left\{ \gamma \left(1 - \frac{V_R}{V_g} \right)^2 \frac{\partial \omega}{\partial t} + \left[\frac{k \lambda v_{\perp}^2}{2 \Omega_e} - \left(1 + \frac{\delta^2}{2} \frac{\Omega_e - \gamma \omega}{\Omega_e - \omega} \right) V_R \right] \frac{\partial \Omega_e}{\partial h} \right\}$$

$$\delta^2 = 1 - \frac{\omega^2}{c^2 k^2}$$

Linear relation (if everything else constant and 1st term dominant)

Making it testable

- Need to eliminate some variables:

$$B_w = F\left(S, \frac{\partial \omega}{\partial t}, k, \omega, V_g, \Omega_e, \frac{\partial \Omega_e}{\partial h}, v_{\perp}, v_{\parallel}, V_R\right)$$

Making it testable

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$$B_w = F\left(S, \frac{\partial \omega}{\partial t}, k, \omega, V_g, \Omega_e, \frac{\partial \Omega_e}{\partial h}, v_{\perp}, v_{\parallel}, V_R\right)$$

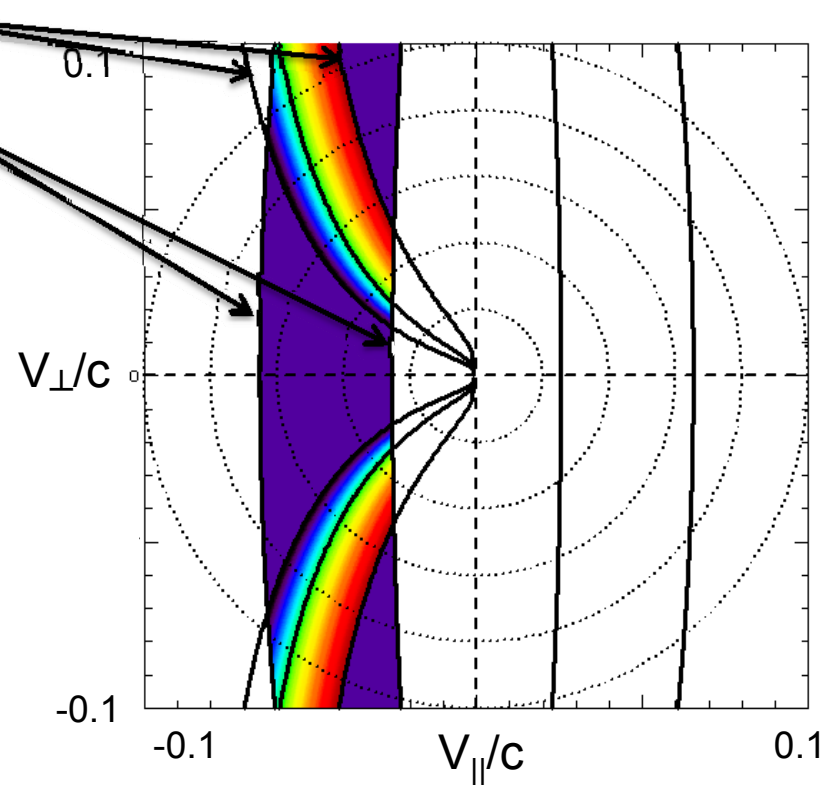
- Growth in range $-0.6 < S < -0.2$
- Dispersion relation (introduces ω_{pe})
- Growth in range $0.2 < f/f_{ce} < 0.5$
- Resonance condition

$$\rightarrow B_w = G\left(\frac{\partial \omega}{\partial t}, \Omega_e, \omega_{pe}, \frac{\partial \Omega_e}{\partial h}, v_{\perp}, v_{\parallel}\right)$$

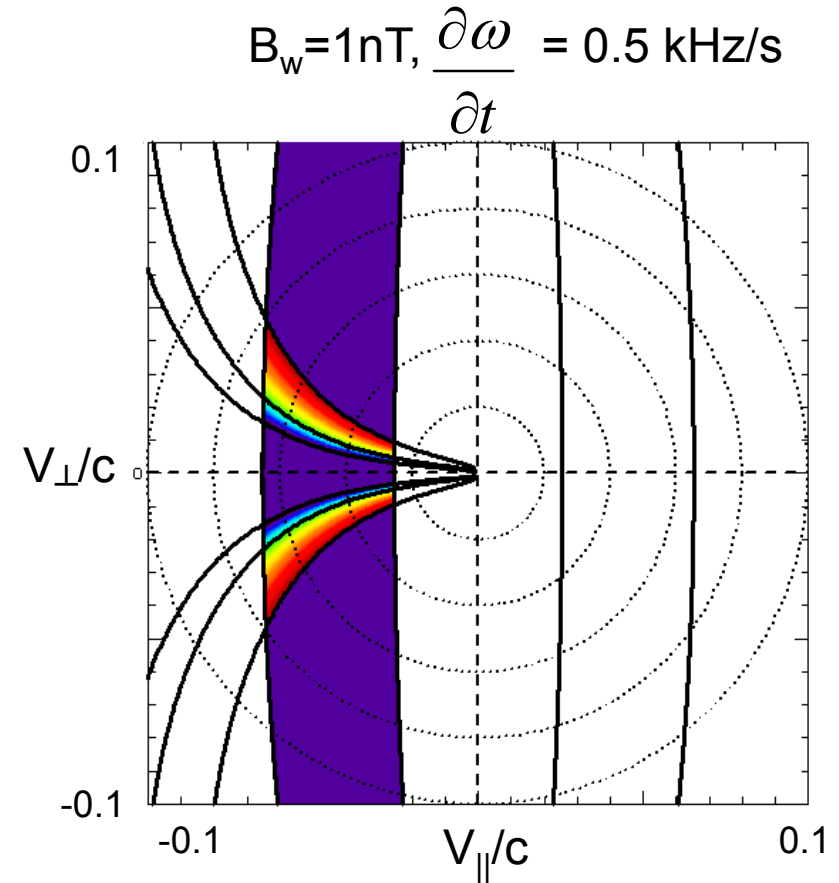
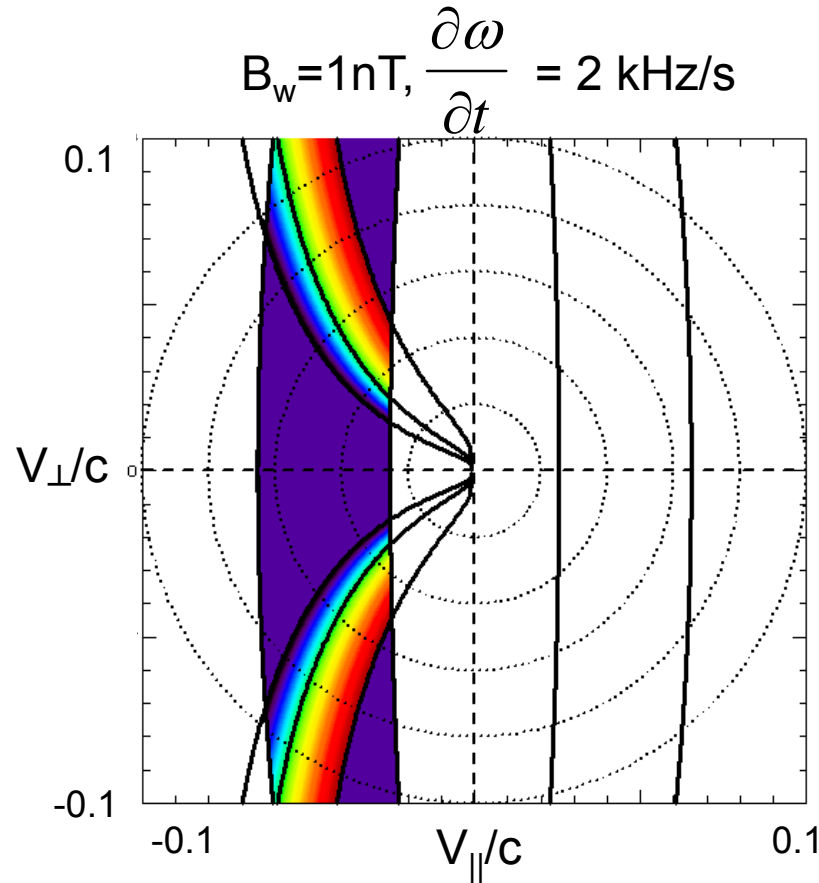
Plasma parameters
Velocities (???)

Maximization in velocity space

- Fastest-growing wave is the one that maximizes the resonant current
 - ★ Maximize $n' = \int f d^3v$ over domain bounded by:
 - $-0.6 < S < -0.2$
 - $0.2 < f/f_{ce} < 0.5$



Maximization in velocity space



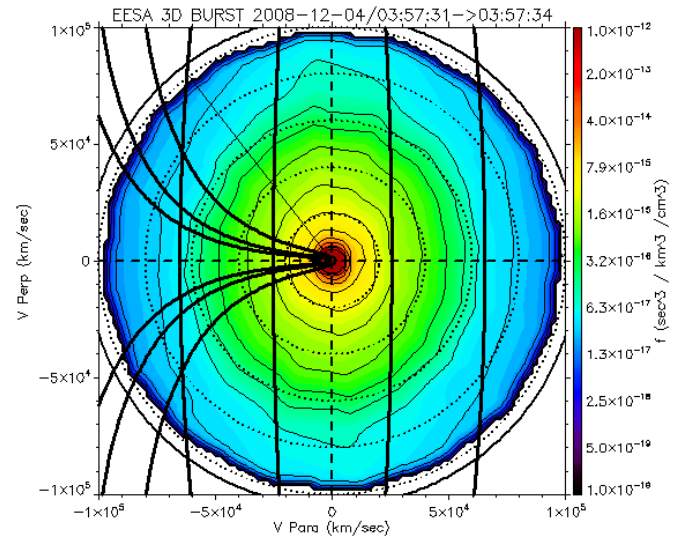
- Integration domain depends on B_w and $\frac{\partial \omega}{\partial t}$ through S

Testable result

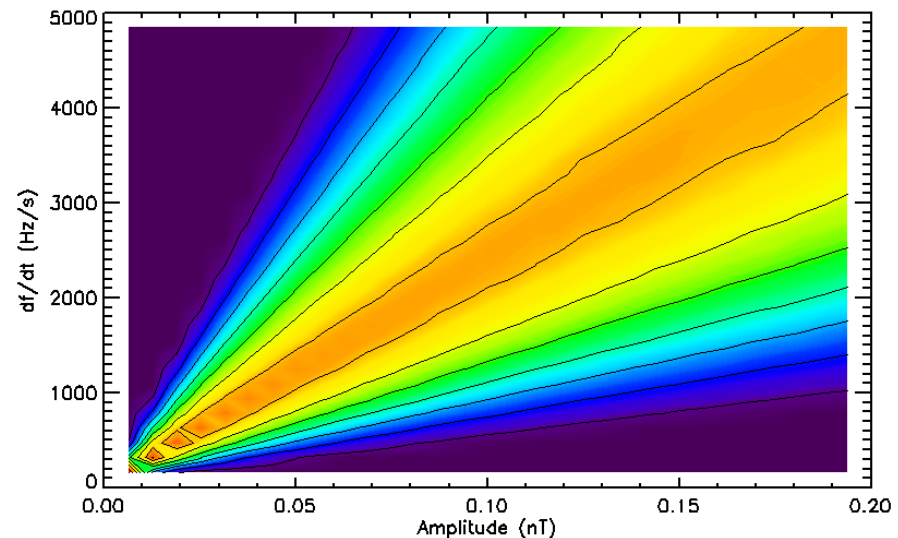
- By integrating over the observed particle distributions, arrive at

$$n' = F\left(\frac{\partial\omega}{\partial t}, B_w\right)$$

- Maximizing n' yields a very specific testable prediction:
 - Relationship between B_w and $\partial f/\partial t$
 - NO adjustable parameters!!

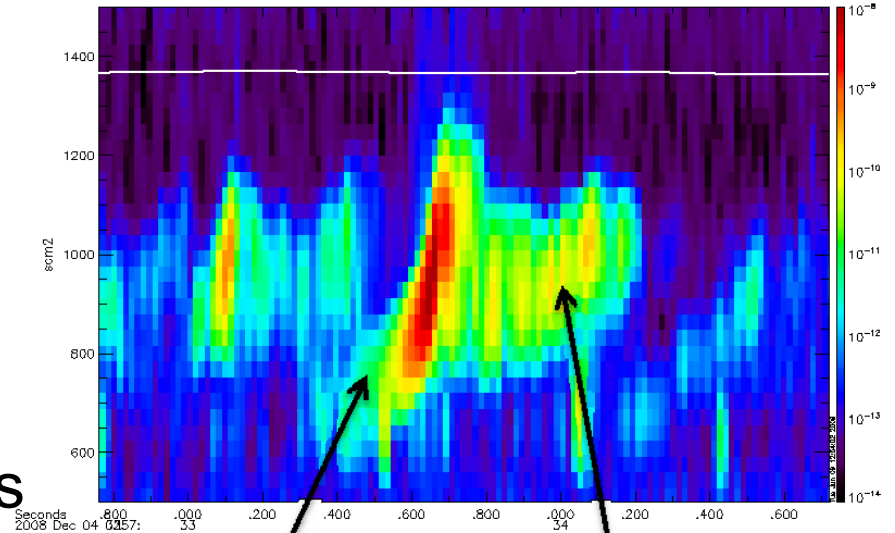


$$n'(B_w, \partial f/\partial t)$$



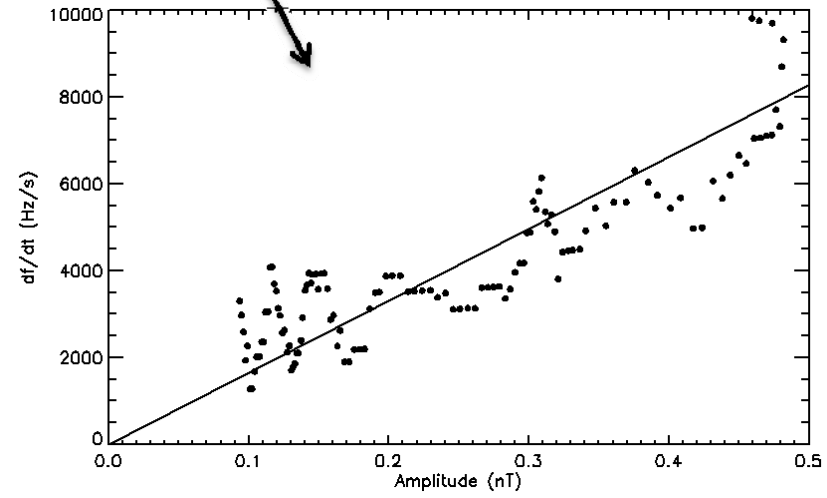
Testing it...

- Need B_w and $\partial f/\partial t$.
Either:
 - ★ Select multiple chorus elements, use average B_w and $\partial f/\partial t$ for each element
 - Amplitude and sweep rates vary throughout sweep
 - ★ Select one element
 - Use zero crossings to get “instantaneous” B_w and frequency



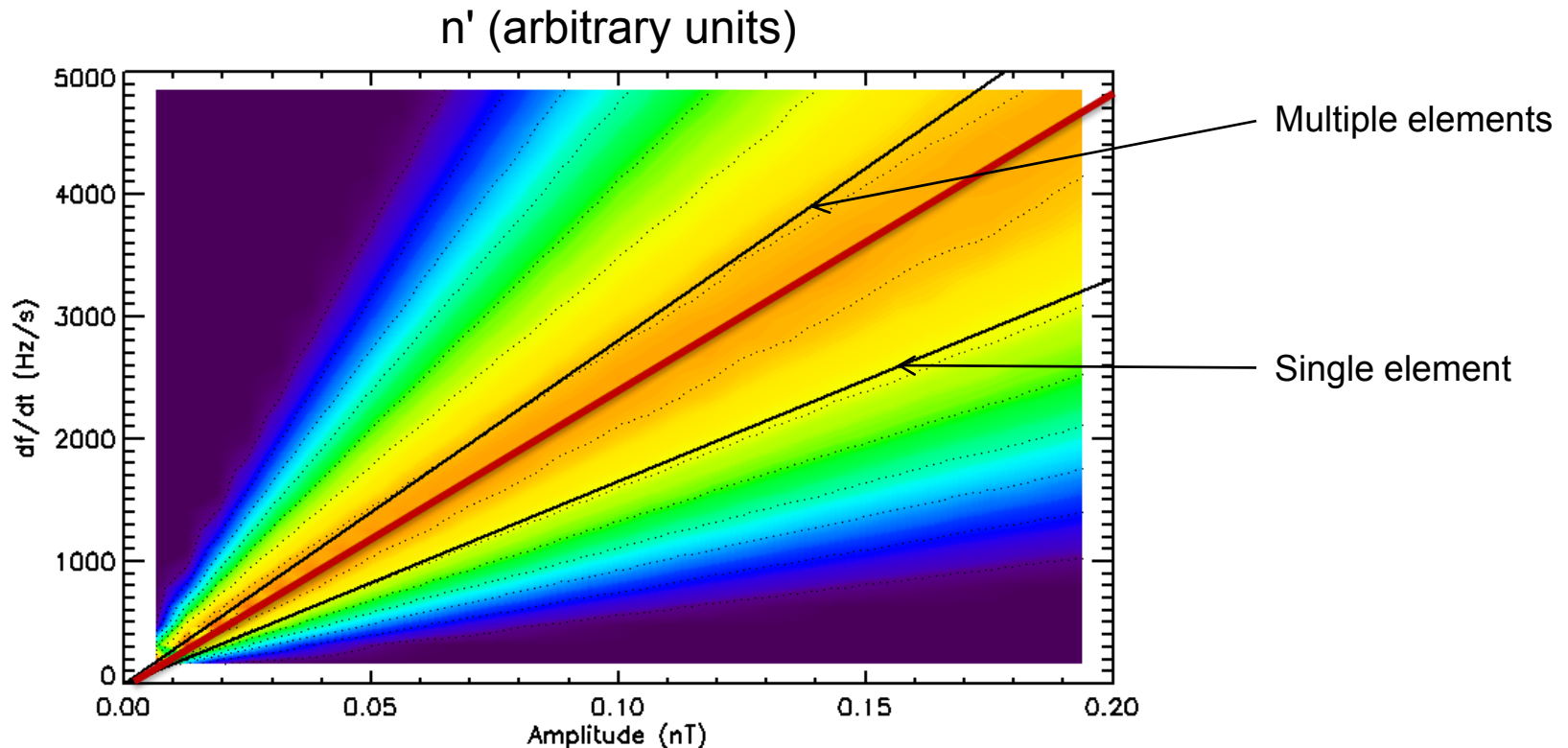
Curved large-amplitude element

Slow small-amplitude element

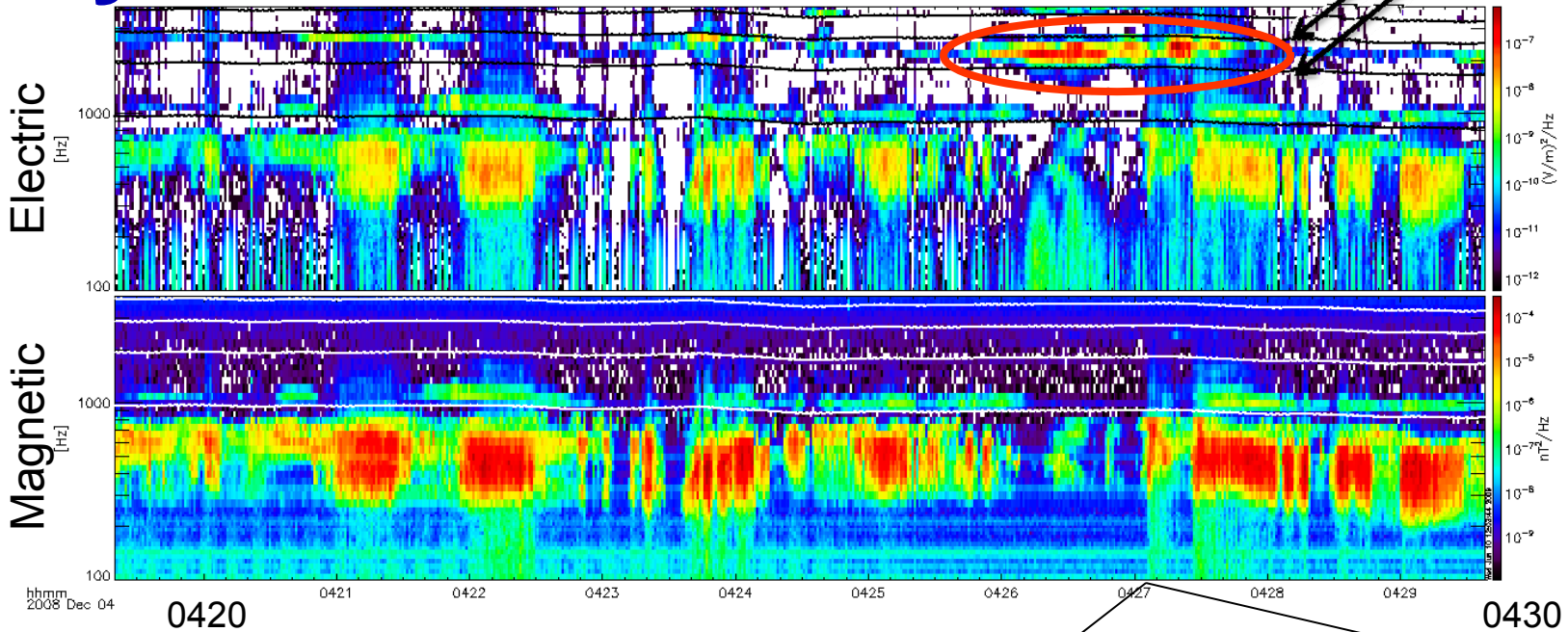


The result

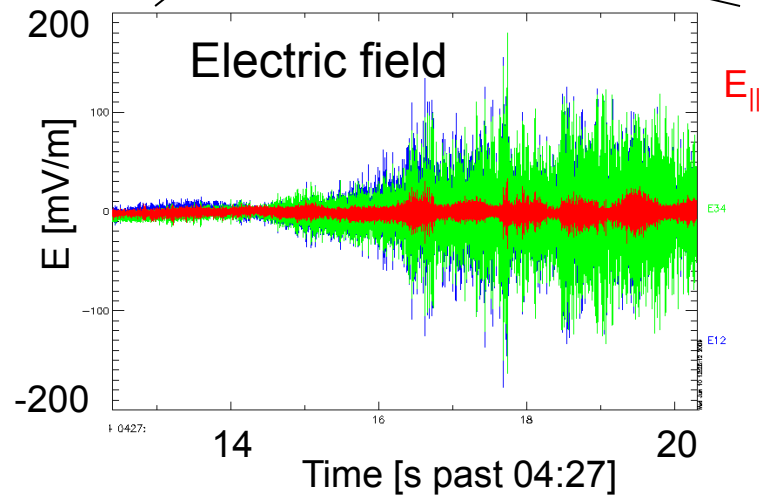
- Observed chorus elements **agree in detail** with the theoretical prediction
 - ✦ Lines should follow maximum in n' (red)



Cyclotron harmonics

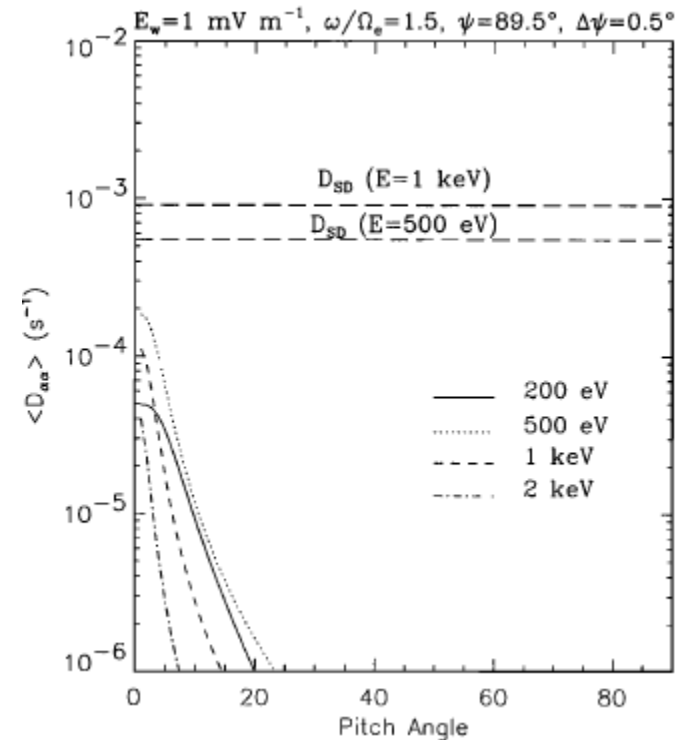


- Electrostatic wave between f_{ce} and $3/2 f_{ce}$
 - ★ Linear polarization
 - ★ Perpendicular to B_0
- Amplitude ~ 100 mV/m
 - ★ Greater than the whistlers
 - ★ Greater than expected



ECH: Scattering rates

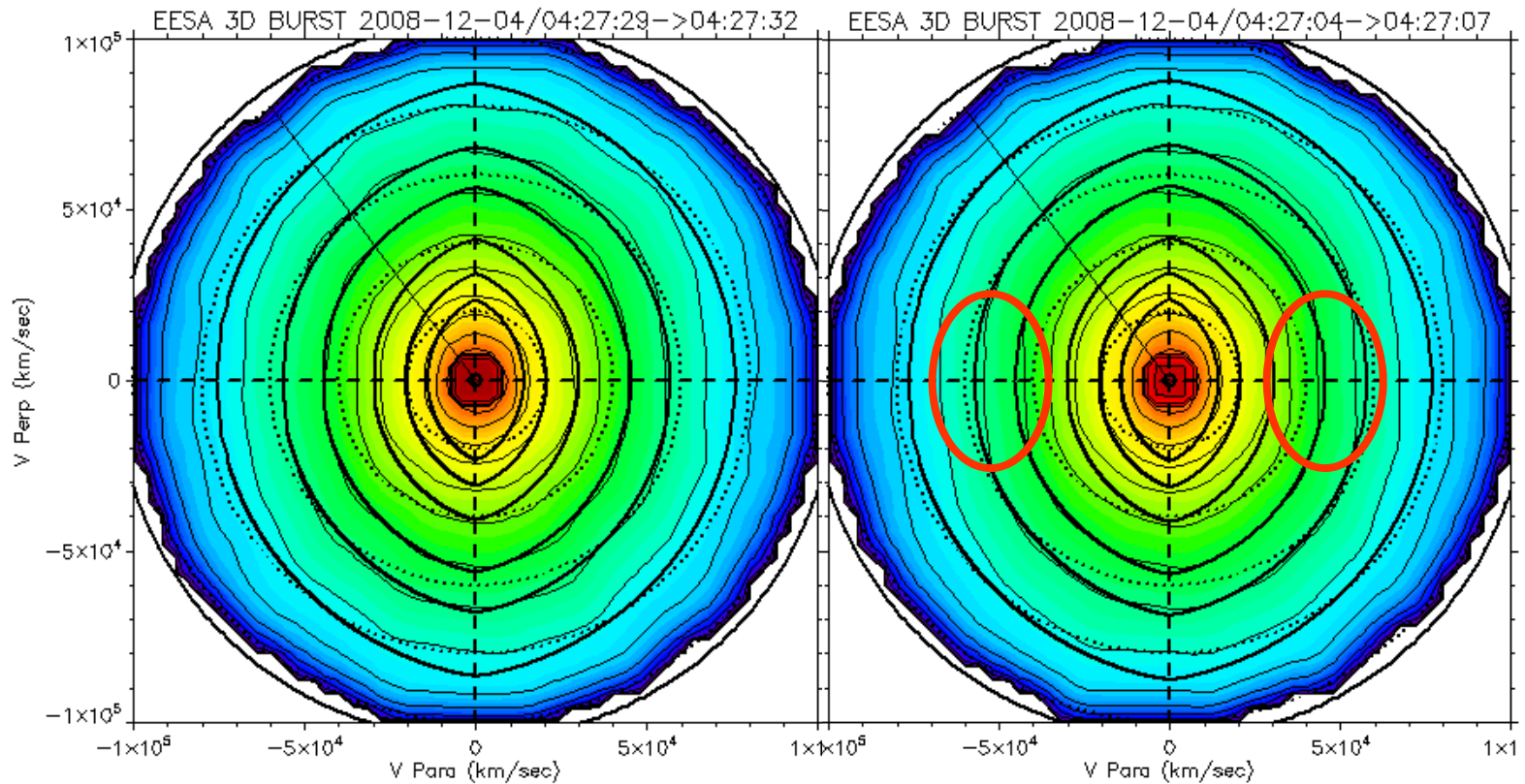
- Pitch angle diffusion rates for Electron Cyclotron Harmonic waves (ECH) calculated by Horne and Thorne.
- For 100 mV/m:
 - ✦ Strong scattering limit over several tens of degrees for ~keV particles
 - Loss cone filled on each bounce



Horne and Thorne, JGR, 2000

ECH: effect

- ECH waves remove “extended” loss cone
- Whistler waves fill it back in



ECH inactive
Whistler active

ECH active
Whistler inactive

Conclusions

- Observation: maximum electric field amplitudes larger than previously seen (>100 mV/m)
- Quasilinear theory:
 - ★ Good fit to diffusion curves
 - ★ Featureless spectrum up to f_{ce} not observed
- Nonlinear theory:
 - ★ Amplitude vs. sweep rate (Omura)
 - Specific testable prediction verified
- Open (?):
 - ★ What absorbs the waves at $f_{ce}/2$?
 - ★ What causes the high-amplitude unstructured oblique whistlers?
 - ★ Interaction between Electron Cyclotron Harmonic waves and chorus whistlers?



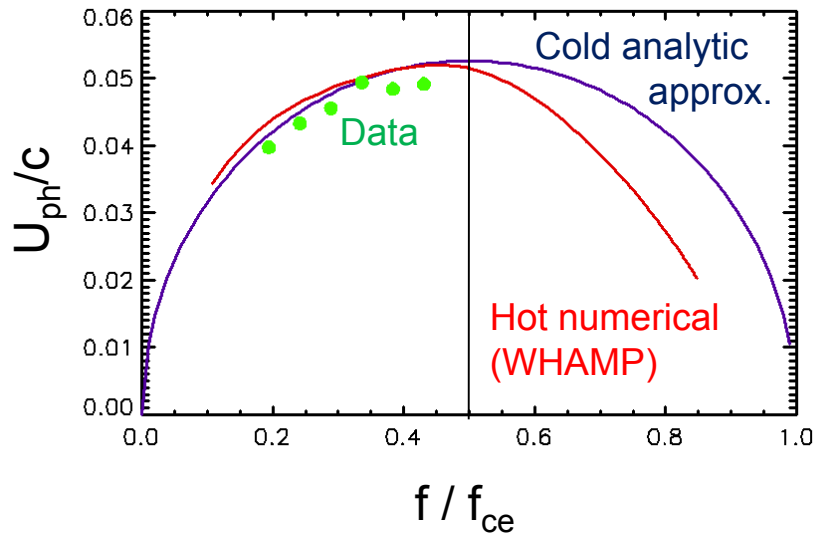
Themis: nonlinear whistlers



EXTRA SLIDES



Dispersion curves



- 2-component plasma: cold isotropic core + hot anisotropic
- Good match to dispersion curve
 - ★ Test of instrument performance
- Thermal effects not important below $0.5 f_{ce}$
 - ★ Wave power confined to range $0.2-0.5 f_{ce}$



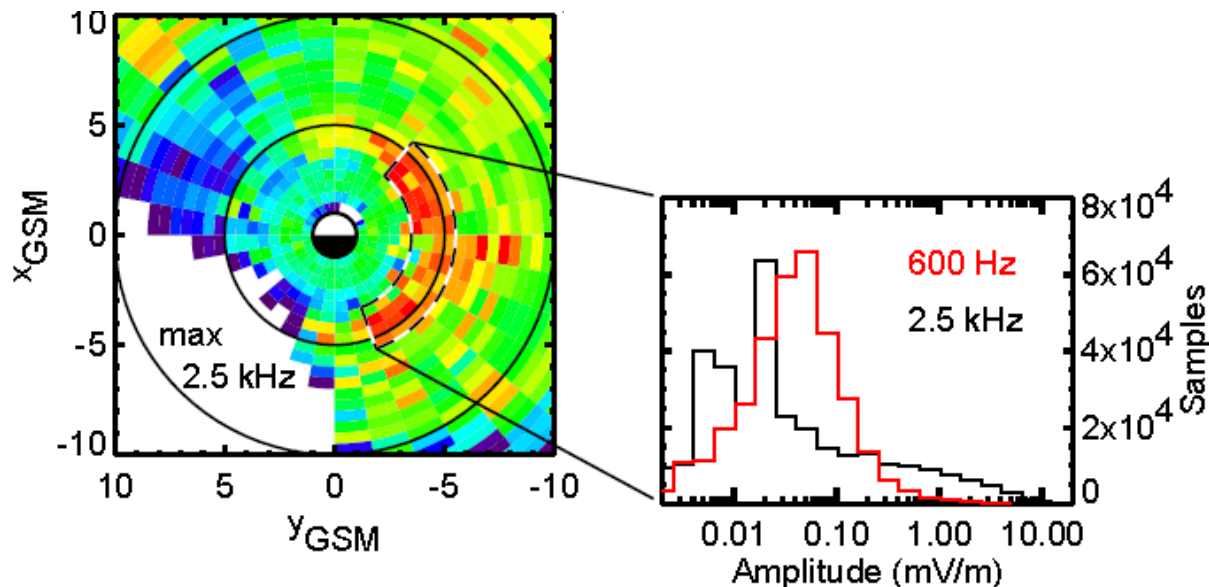
Themis: nonlinear whistlers

STATISTICAL RESULTS

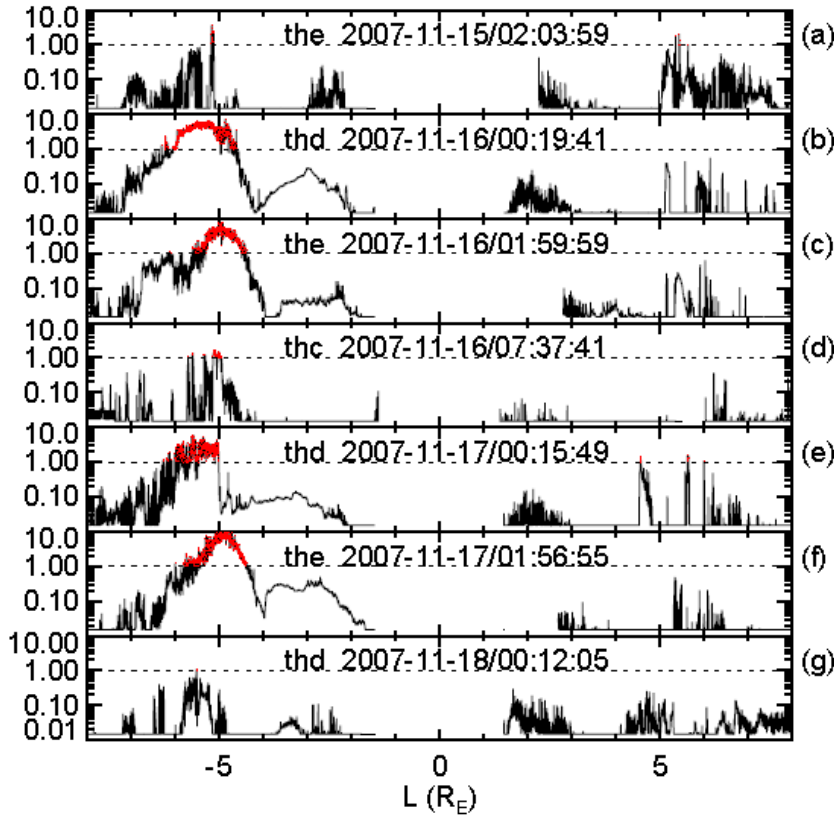


Electric field amplitude distribution

- Probability distribution function is "heavy"-tailed, with substantial probability out to tens of mV/m.
 - ✿ 4-s averaged data: values up to tens of mV/m
 - Instantaneous values much larger, but limited statistics



Spatial/temporal distribution



- Multiple satellites yield information on spatial and temporal distribution
- Enhanced activity can last for days
- Confined to a few hours of local time

Statistics: conclusions

- Whistler electric field amplitudes have a heavier tail than previously reported
 - ★ > 100 mV/m bursts
 - ★ High-amplitude regions are:
 - Spatially extended (several R_E)
 - Temporally persistent (several days)
 - ★ Relevant for acceleration/loss processes
- Many of the highest-amplitude events are obliquely propagating
 - ★ Even at magnetic latitudes < few degrees
 - ★ Lack the characteristic chorus chirp